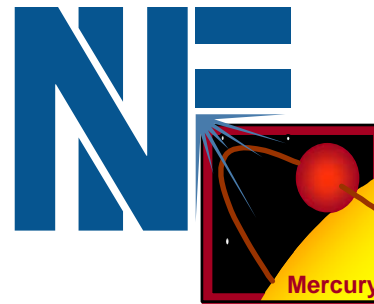


# Status of the Mercury Laser

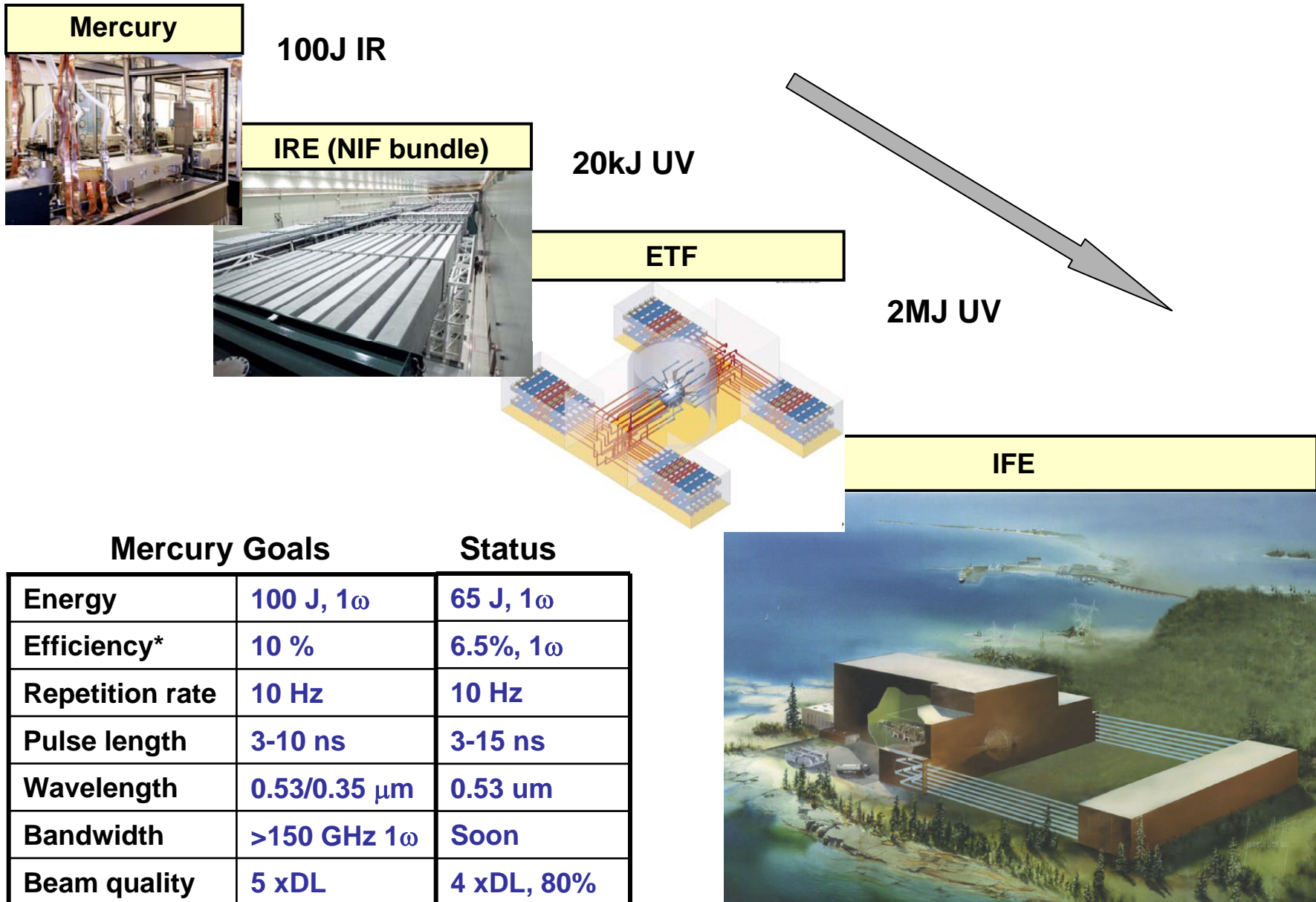
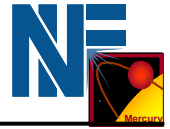


**Andy Bayramian**

**National Ignition Facility Directorate  
Lawrence Livermore National Laboratory  
Livermore, California 94550**

**High Average Power Laser  
Program Workshop  
Rochester, NY  
November 8-9, 2005**

# Mercury is the first step toward a 10 Hz class of IFE lasers



\*utilities not included



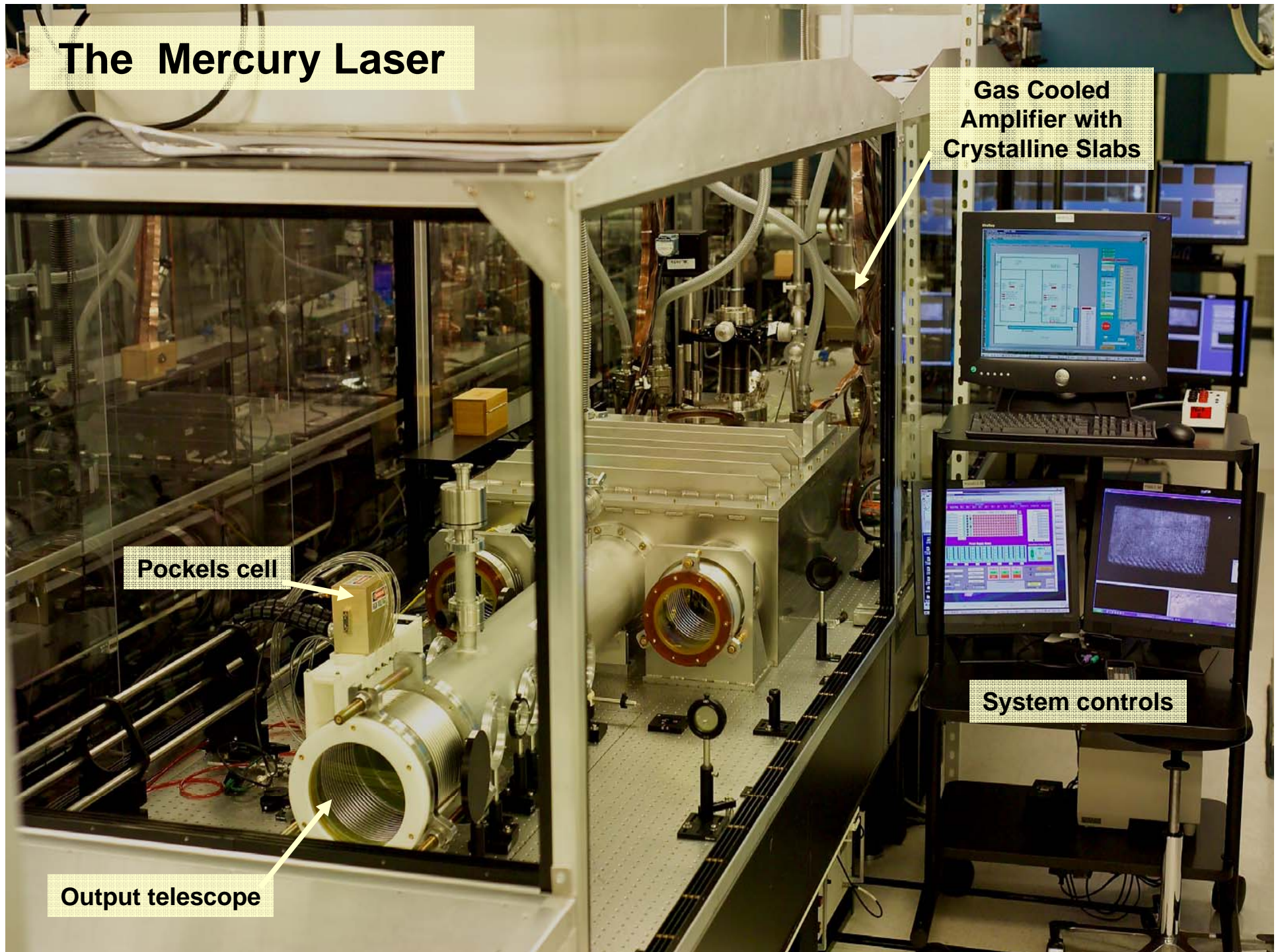
# The Mercury Laser

Gas Cooled  
Amplifier with  
Crystalline Slabs

Pockels cell

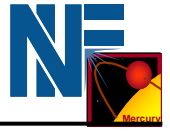
Output telescope

System controls





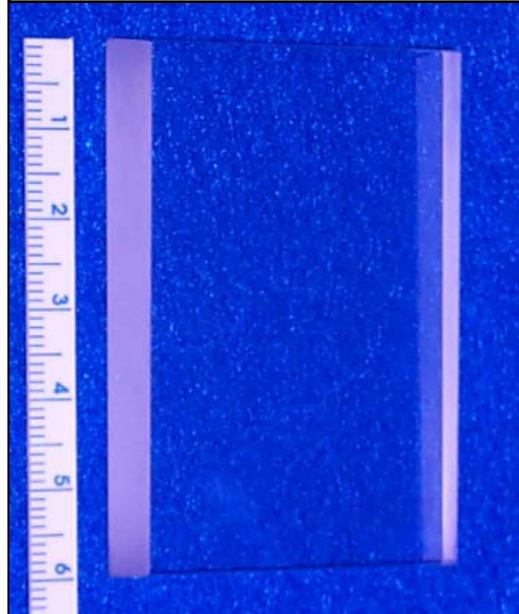
# The Mercury Laser technologies



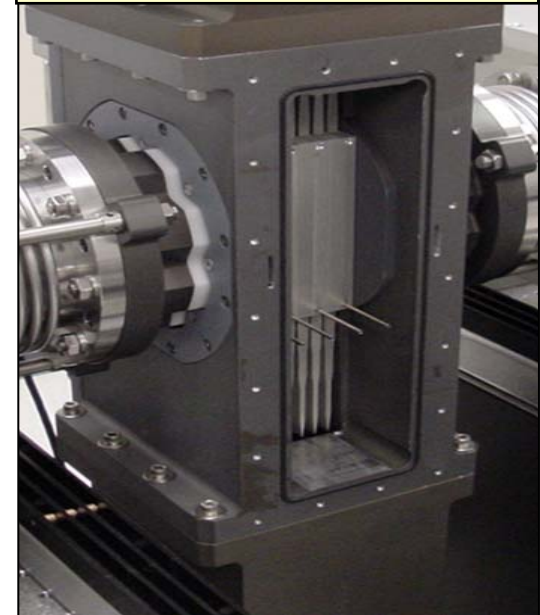
**Diode pump arrays**



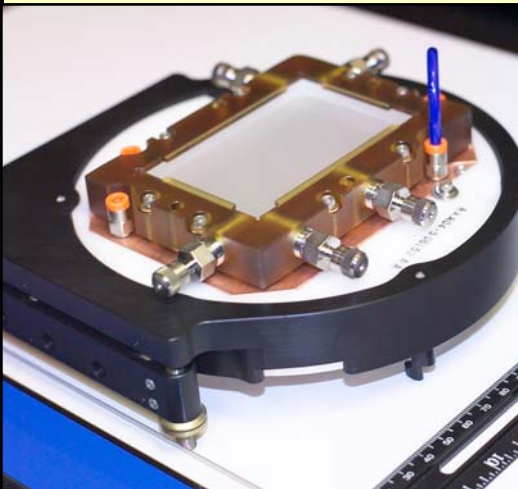
**Solid-state amplifier**



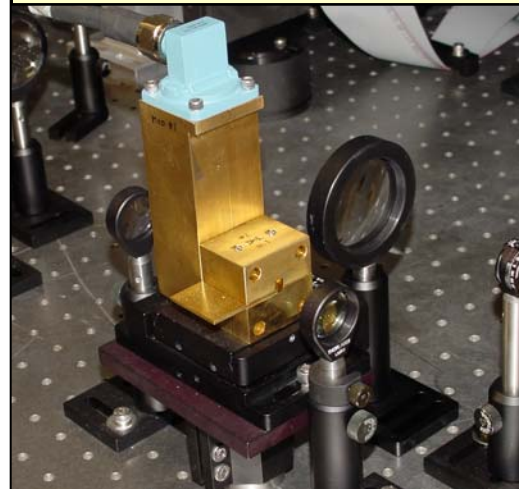
**Helium gas cooling**



**Frequency Converter**



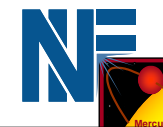
**Broadband Front End**



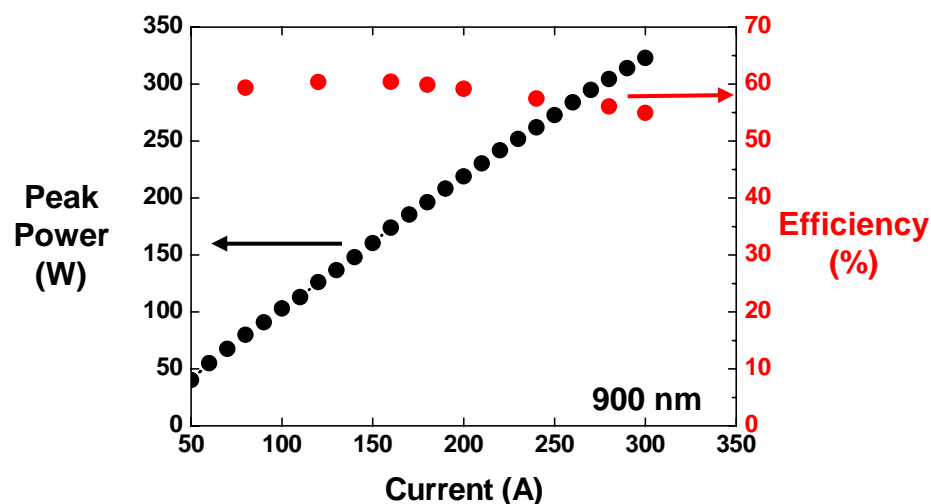
**Adaptive Optic**



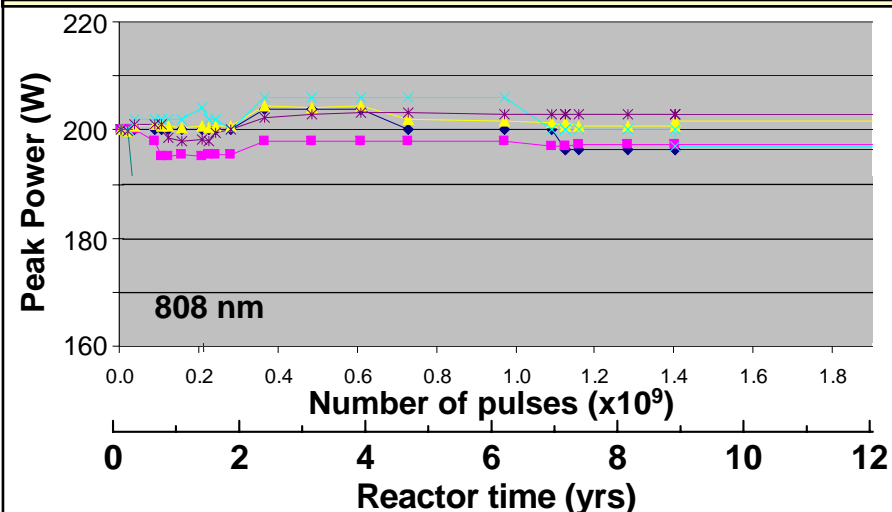
# New diode technology demonstrates enhanced reliability, efficiency, and lower cost per Watt



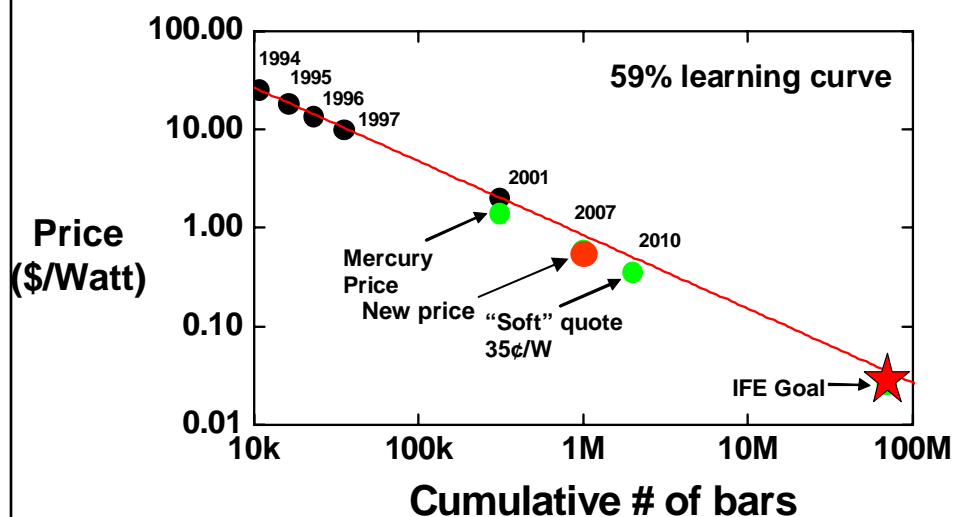
## 325W from 'Superbar' exceeds 300W goal



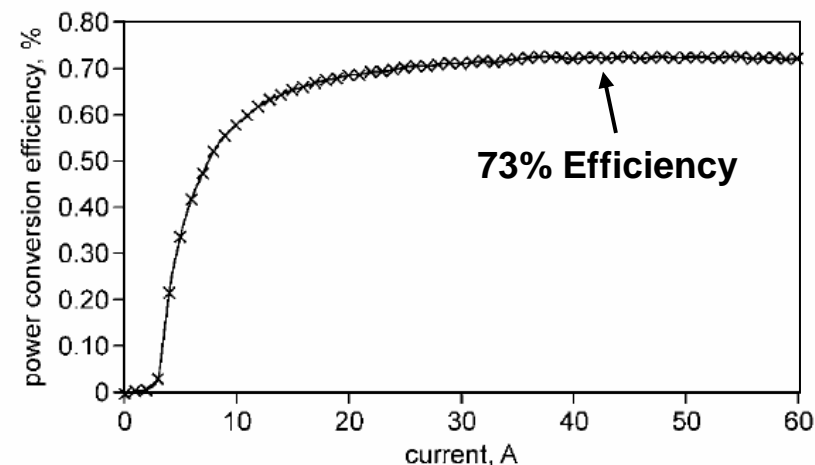
## 'Superbar' lifetime near $10^{10}$ shot goal



## New price point falls in the learning curve

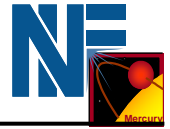


## New efficiency data exceeds 70% goal

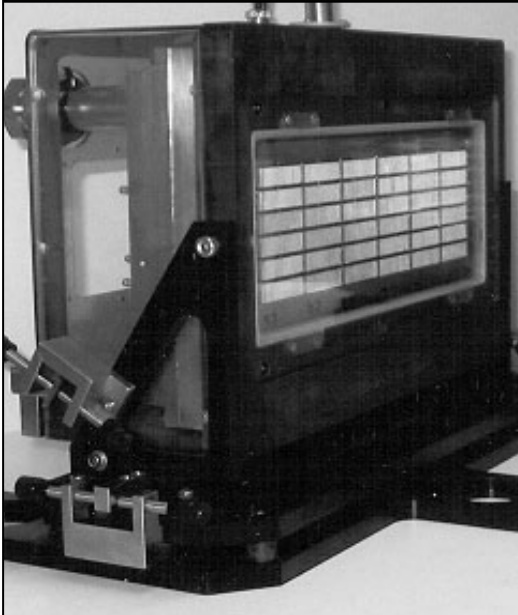


Kanskar, M. et. al. "73% CW power conversion efficiency at 50W from 970 nm diode laser bars", *Elec. Lett.*, 41(5) 2005.

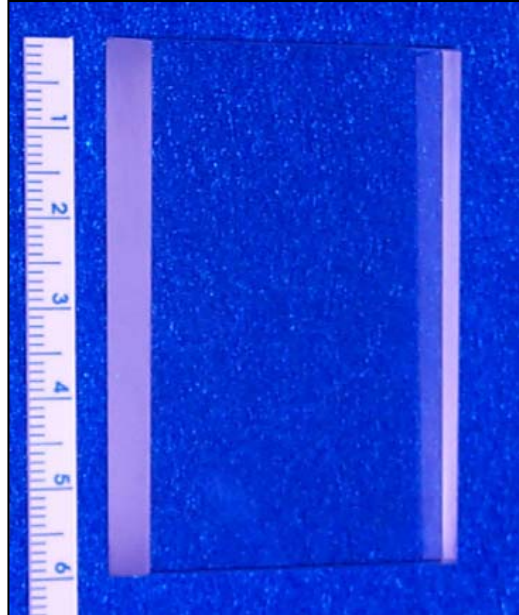
# Commercial Yb:S-FAP growth efforts are currently focused on scaling to kJ apertures



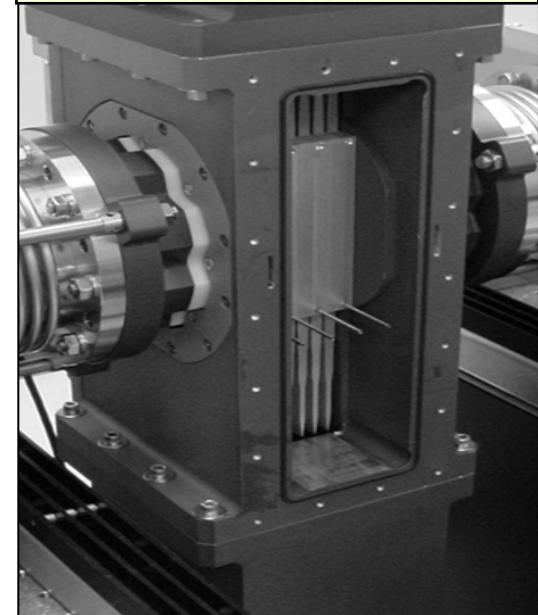
**Diode pump arrays**



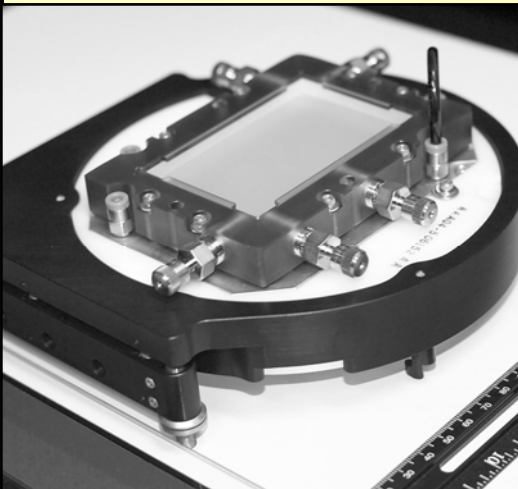
**Solid-state amplifier**



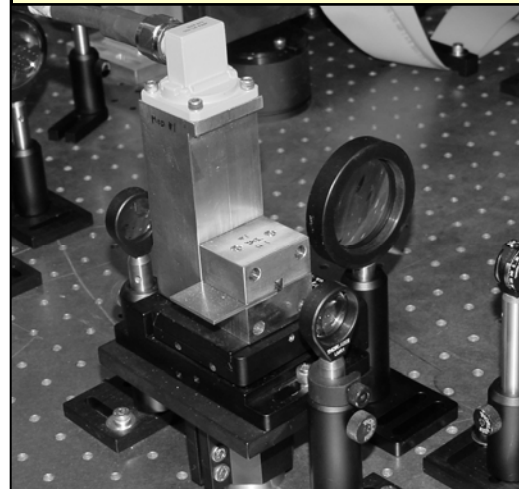
**Helium gas cooling**



**Frequency Converter**



**Broadband Front End**

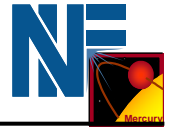


**Adaptive Optic**





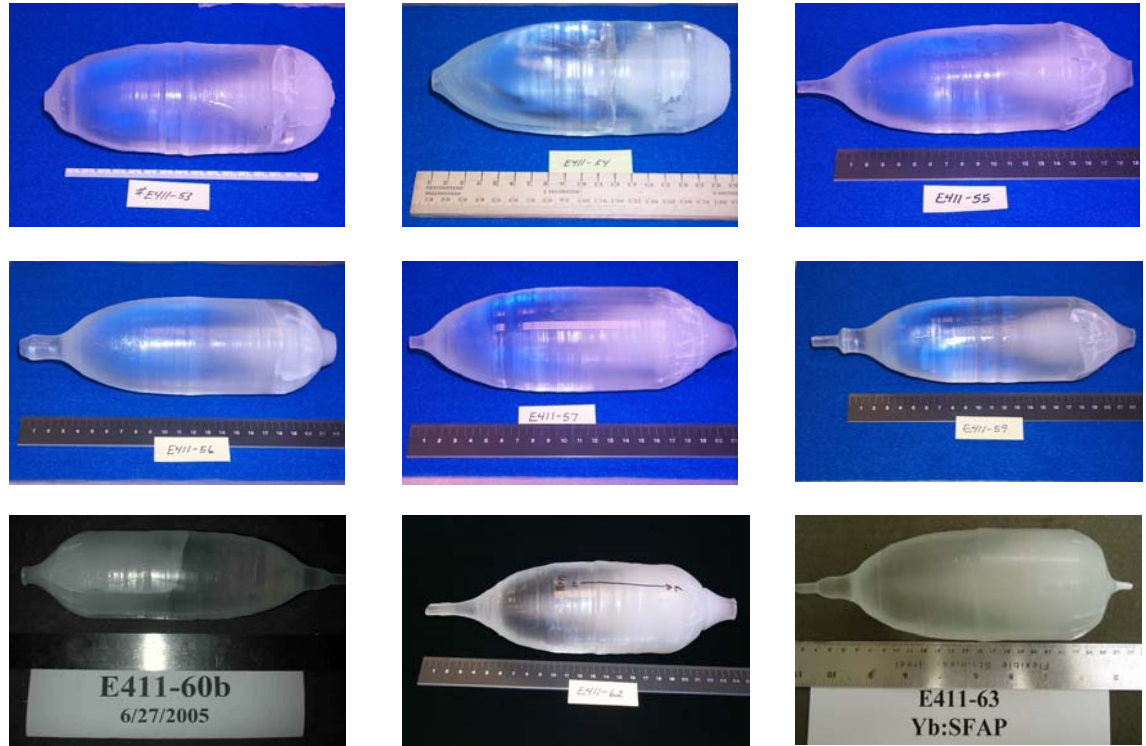
# Successful growth of Mercury-scale boules will lead to kJ-scale boules in 2006



## Northrop Grumman Czochralski growth station

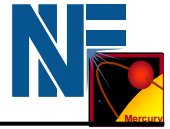


## Boules grown in FY05 which yielded slabs

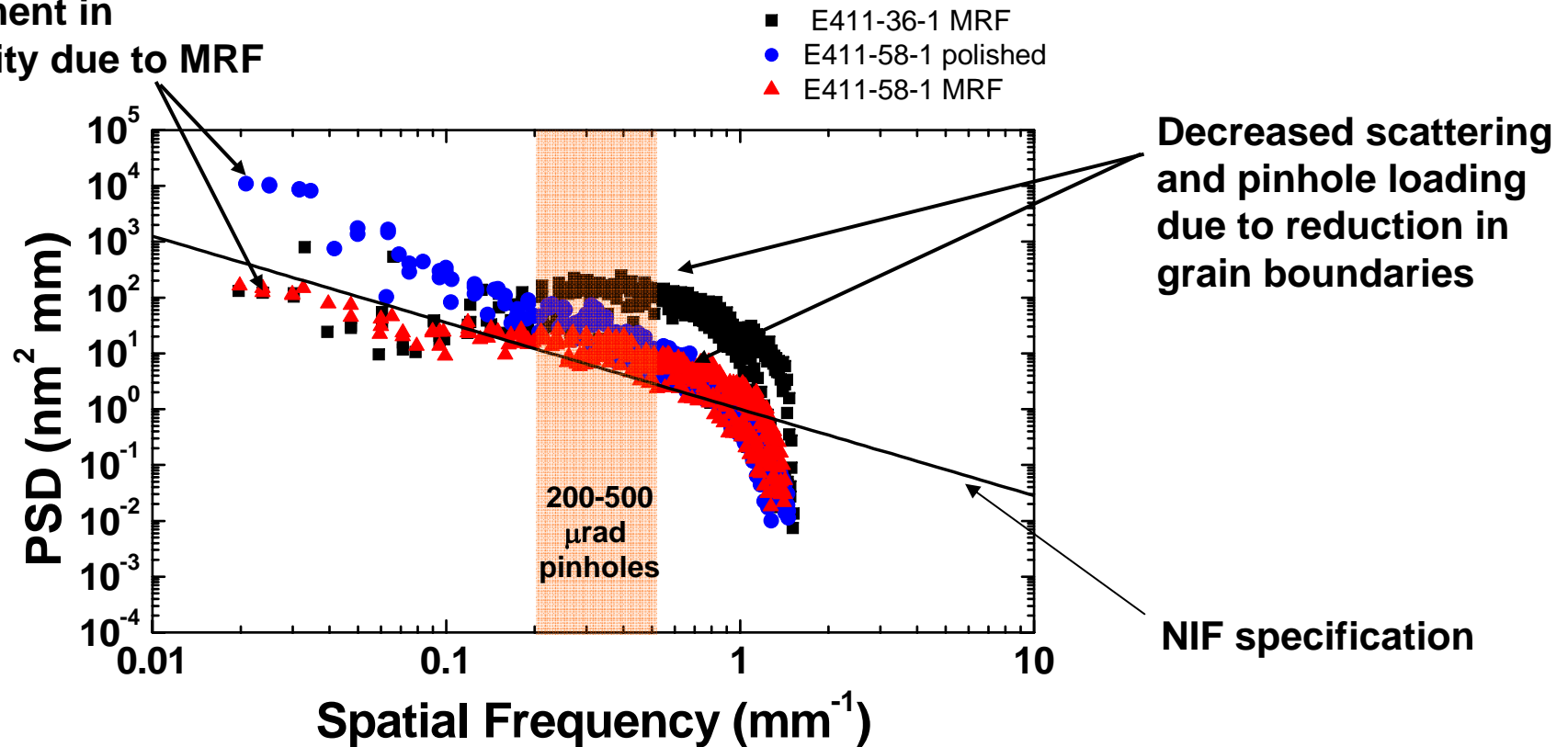


- 23 spare slabs are now in commercial production
- First 10 cm diameter growth in early CY2006
- 10 cm diameter boules can produce 10 x 15 cm single crystal apertures or bonded 20 x 30 cm apertures capable of 1 - 4 kJ

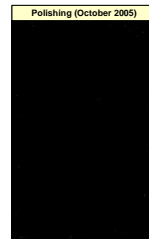
# Optical quality of S-FAP has improved through reduction of grain boundaries, improved polishing, and magneto-rheological finishing



Improvement in focusability due to MRF



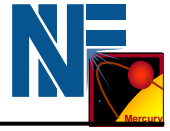
Improved polishing technique increases the laser damage threshold



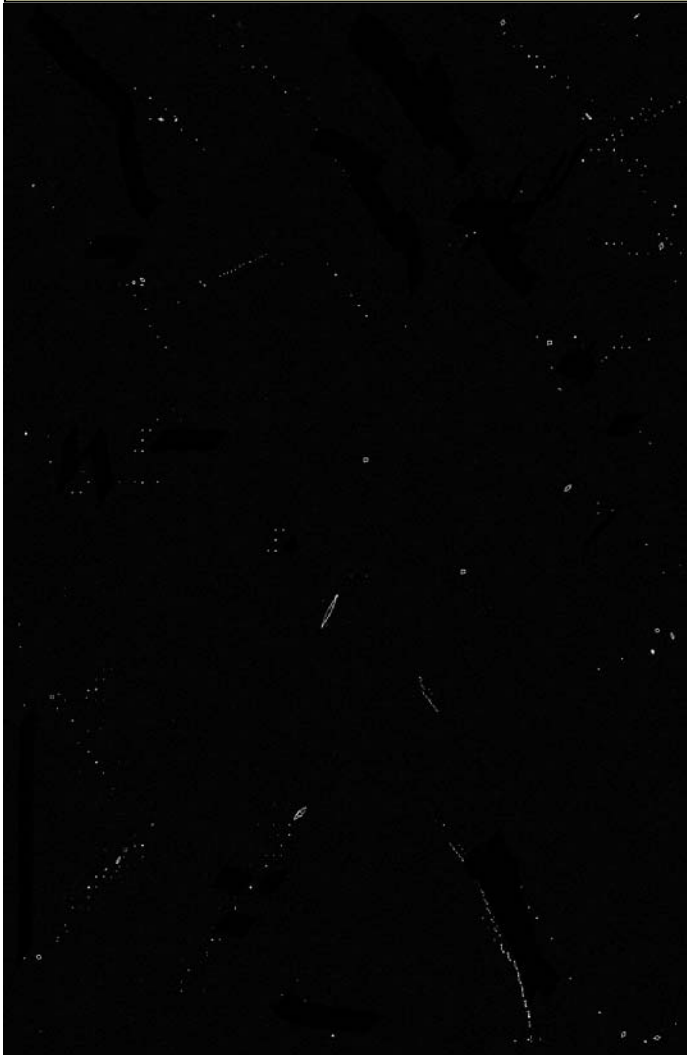


**Optical quality of S-FAP has improved through reduction of grain boundaries, improved polishing, and magneto-rheological finishing**

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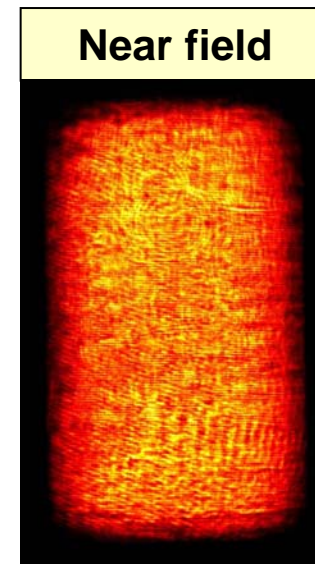
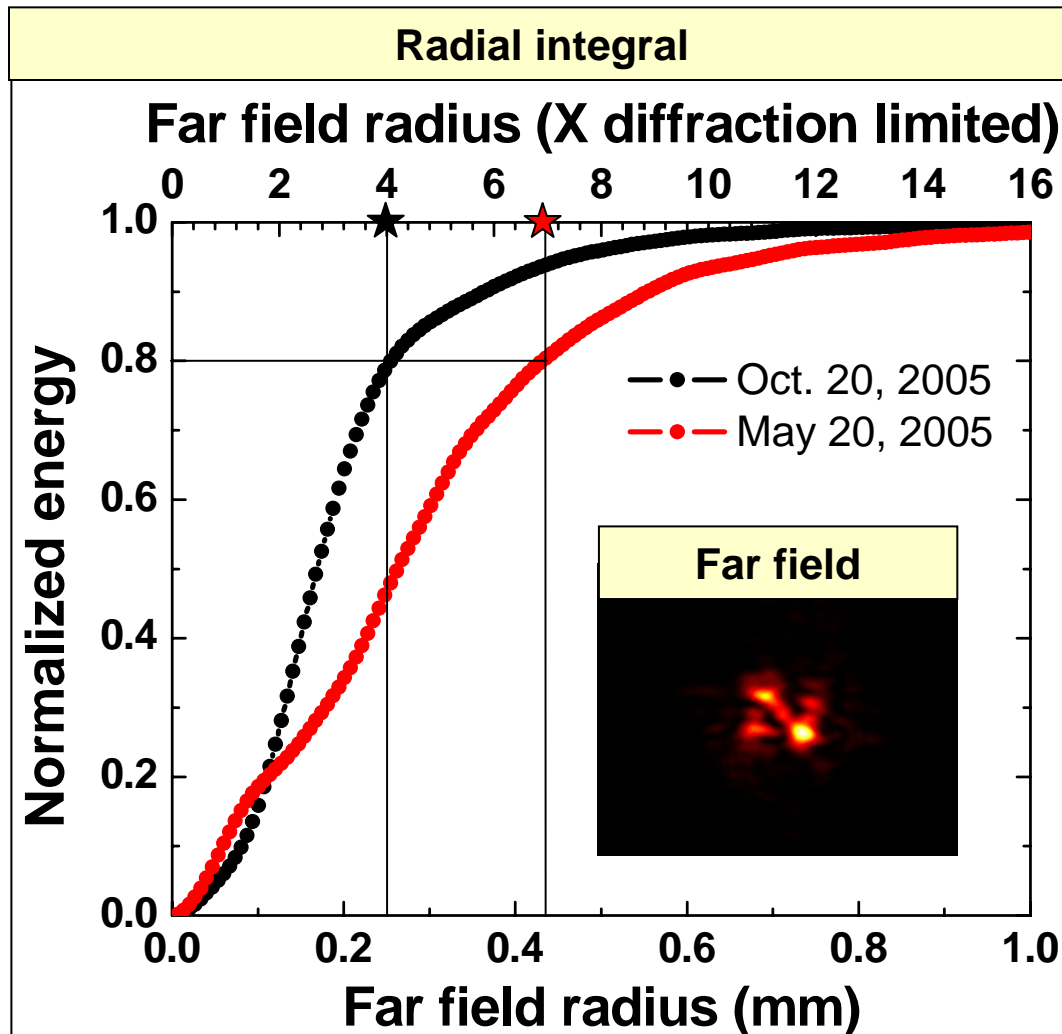
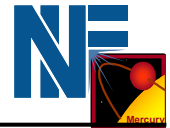
**Polishing (June 2005)**



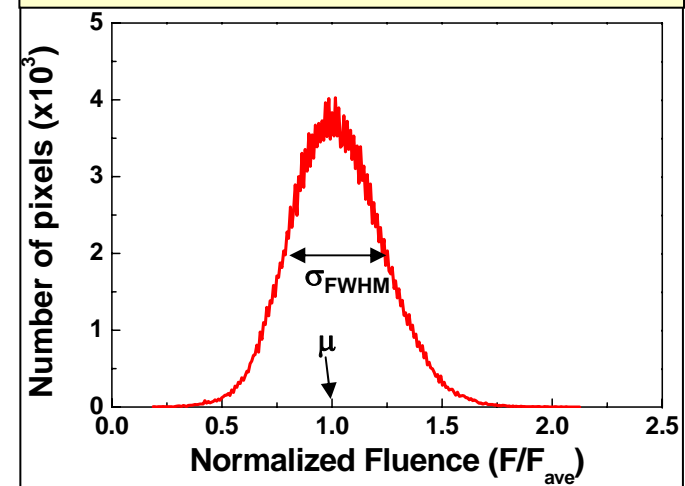
**Polishing (October 2005)**



Far field analysis indicates that we will meet the  
< 5X diffraction limited goal

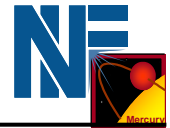


Experimental contrast ( $\sigma/\mu$ ) = 0.44  
Improvements will appear in FY06

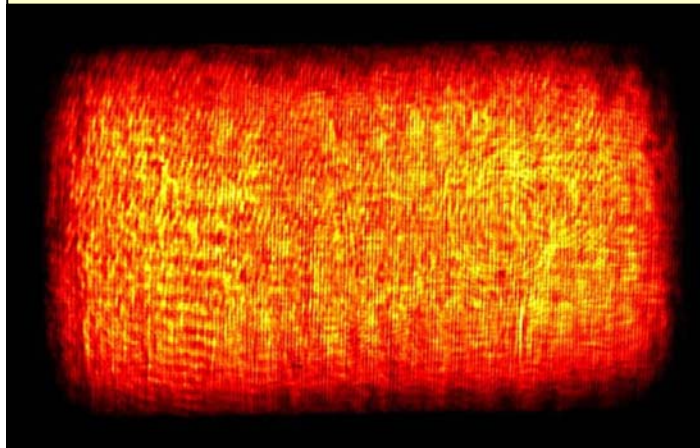


Improvements with new slabs and the adaptive optic

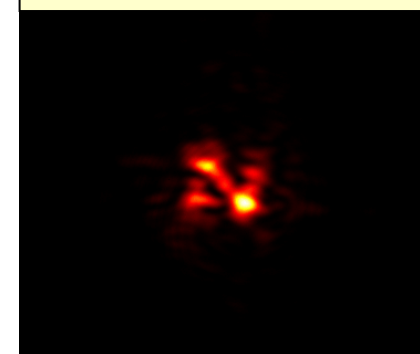
# Output beam stability has been measured



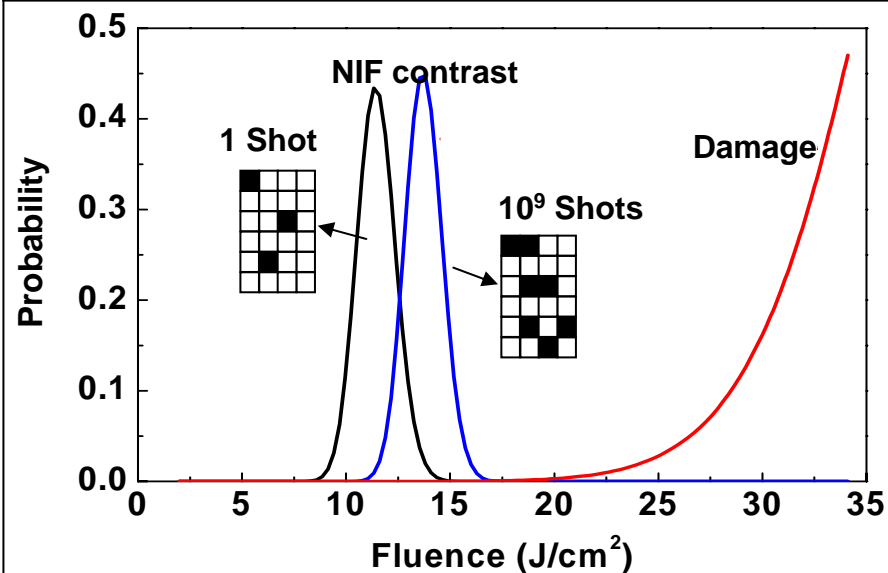
Output Near Field



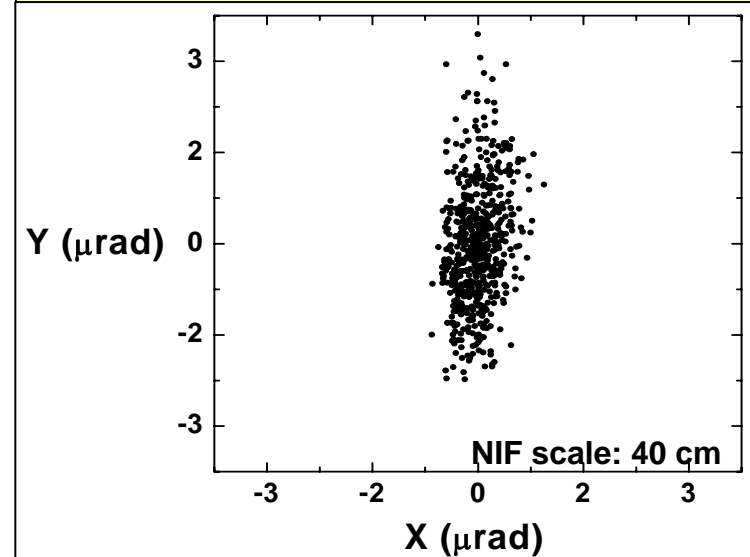
Output Far Field



Max of N analysis indicates small change in fluence for a large numbers of shots

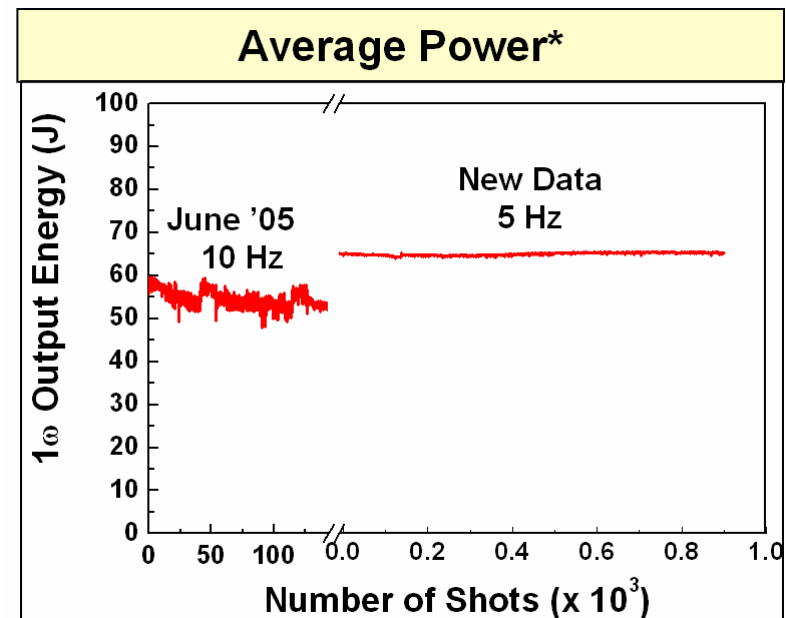
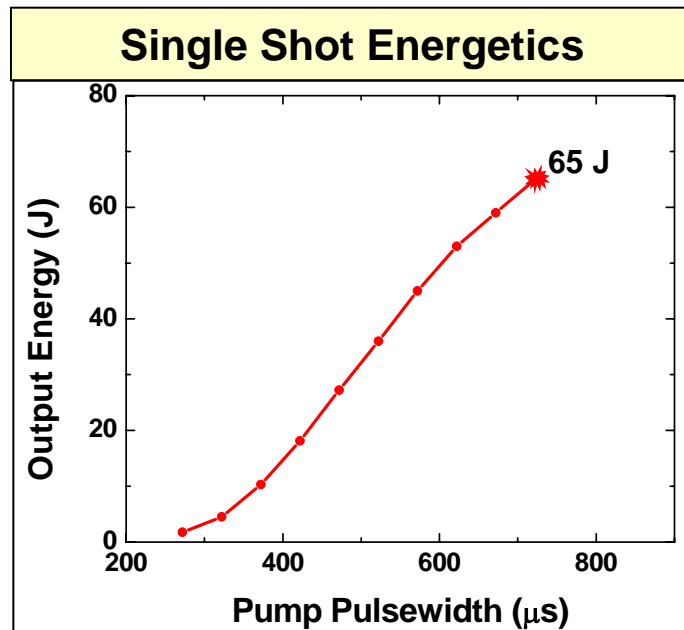
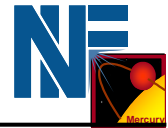


Pointing analysis indicates 0.65  $\mu$ rad RMS stability



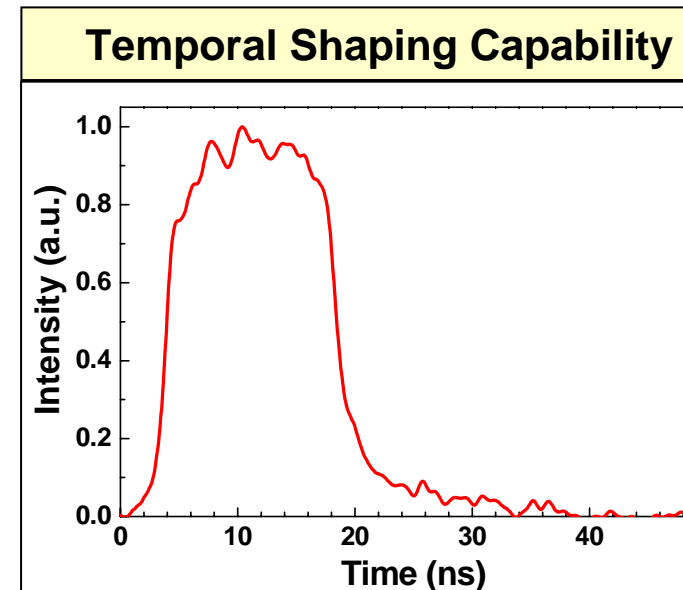


# System upgrades have improved stability and reliability for high power operation



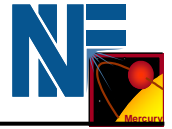
## System upgrades

- Modified multiplexing and U-turn layout to reduce fluence on turning mirrors
- Improved shot and diagnostic stability through better shielding and electronics
- Automated alignment process for active correction of thermal aberrations and long term drift

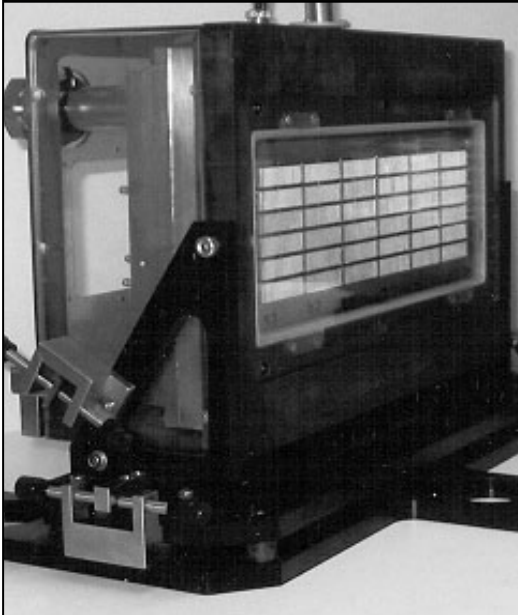


\* Front end drop outs not shown

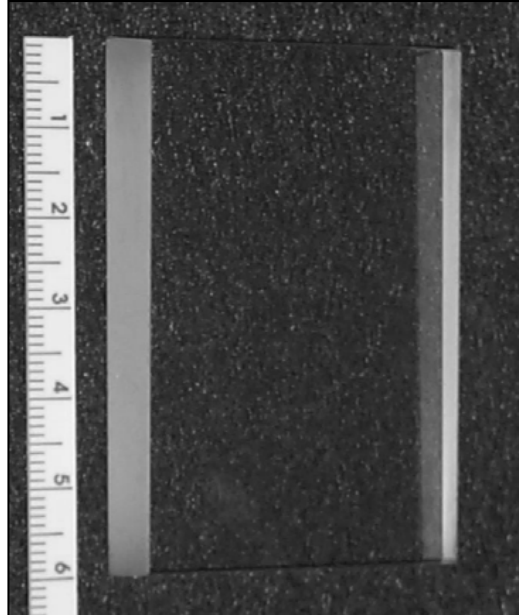
**YCOB has generated over 220 W as an average power frequency conversion material**



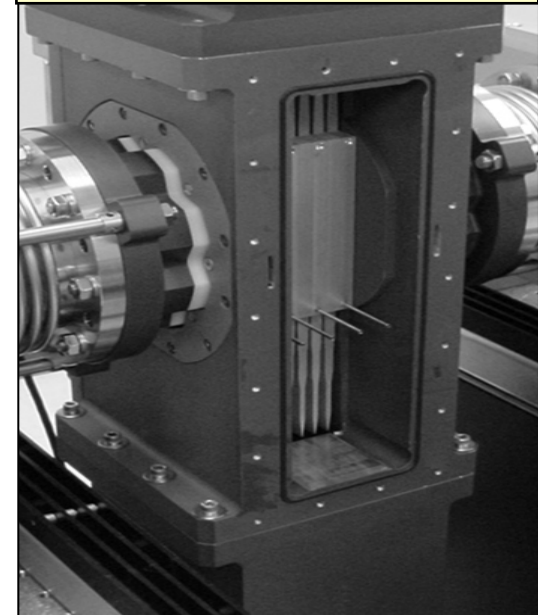
**Diode pump arrays**



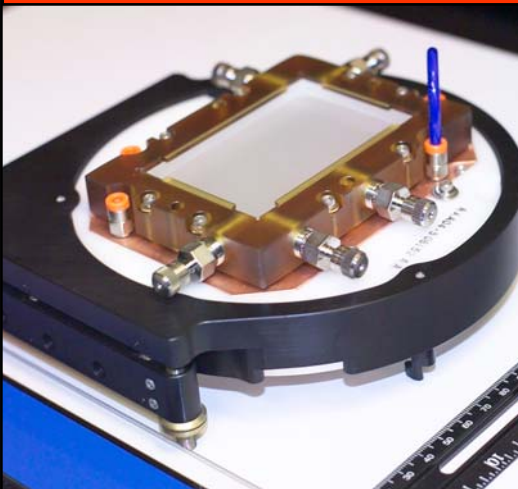
**Solid-state amplifier**



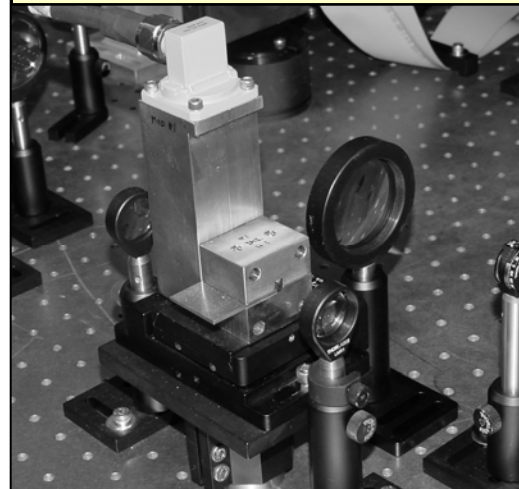
**Helium gas cooling**



**Frequency Converter**



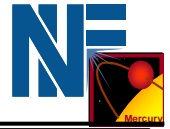
**Broadband Front End**



**Adaptive Optic**

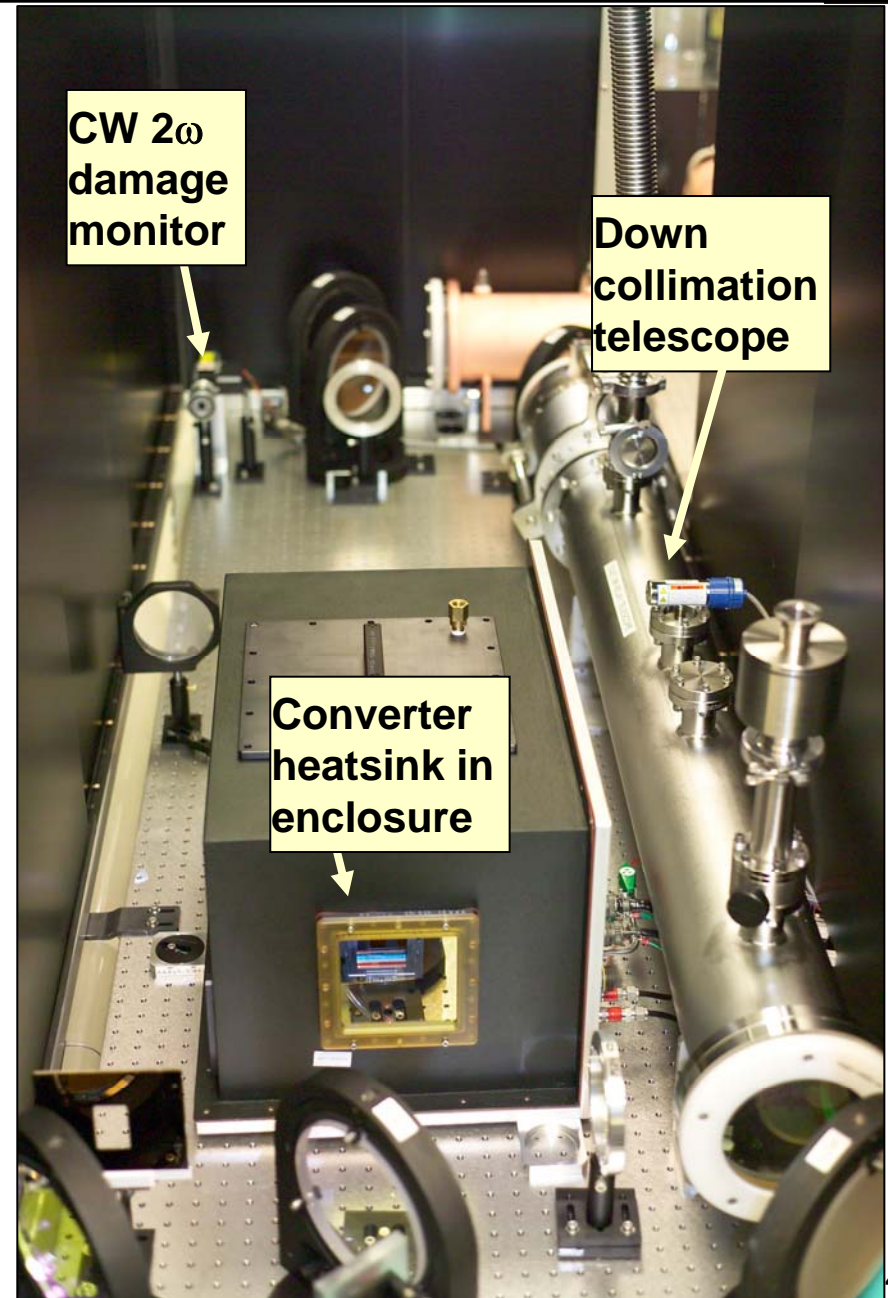
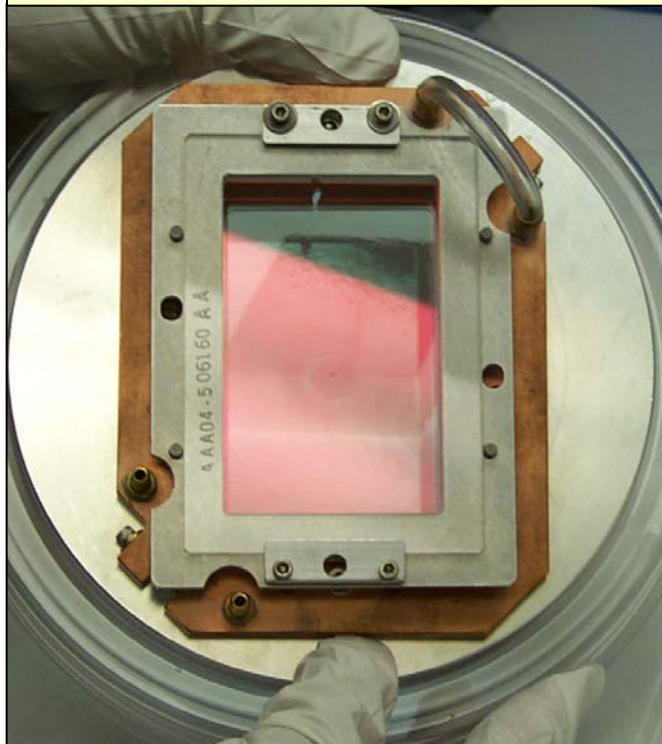


## We have fabricated three full aperture YCOB frequency converter crystals



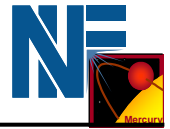
- Aperture size is  $5.5 \times 8.5 \times 1.1 \text{ cm}^3$
- Boules are production growth from Crystal Photonics Inc.
- Czochralski growth method is scalable to kJ apertures

YCOB Crystal in heatsink





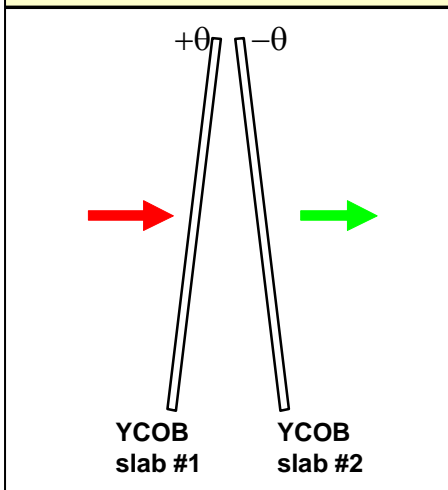
# Multi-plate YCOB frequency converter is efficient up to a bandwidth of 384 GHz ( $1\omega$ ) meeting SSD requirement



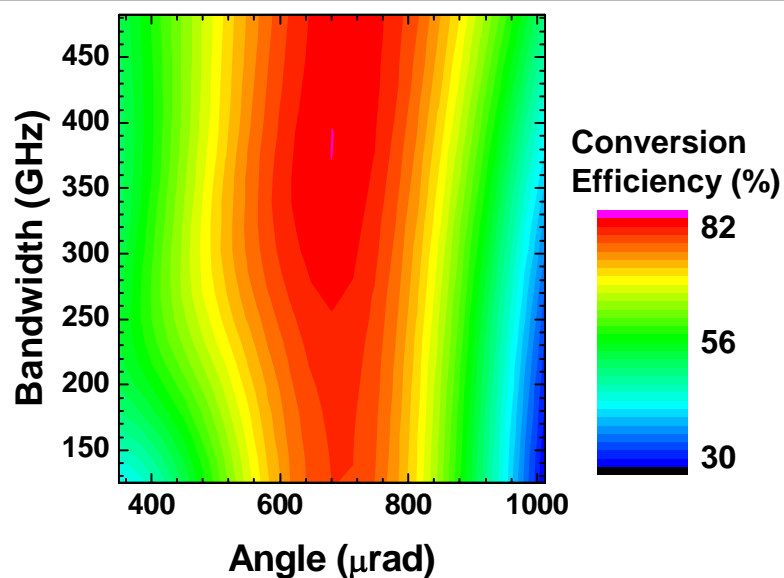
**Goal :** Determine the largest bandwidth and the best phase matching angle for two 7.5 mm thick YCOB crystals

**Result :** Parametric studies give  
 $\Delta\nu = 384$  GHz,  $\theta = 700$   $\mu$ rad

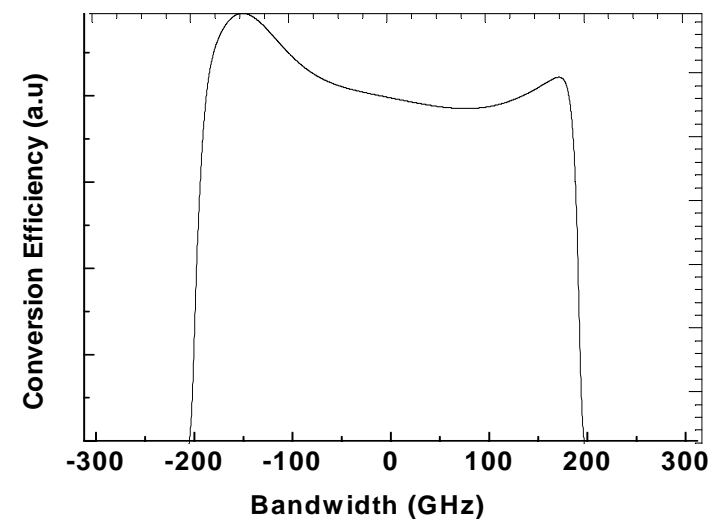
## Multi-plate geometry



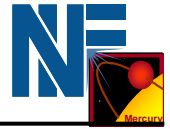
## Efficiency optimization vs bandwidth



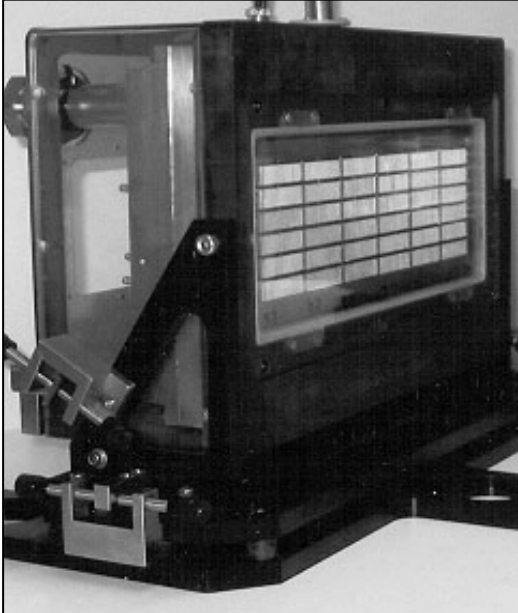
## $2\omega$ spectrum at 768 GHz



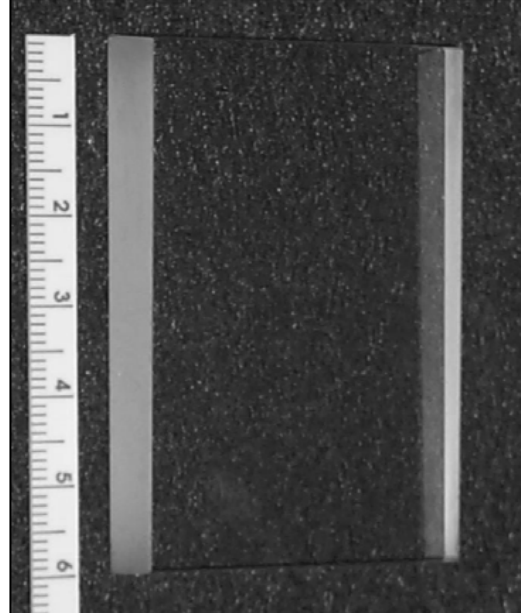
# A temporally, spectrally, and spatially sculpted front end will allow efficient laser extraction and coupling to fusion targets



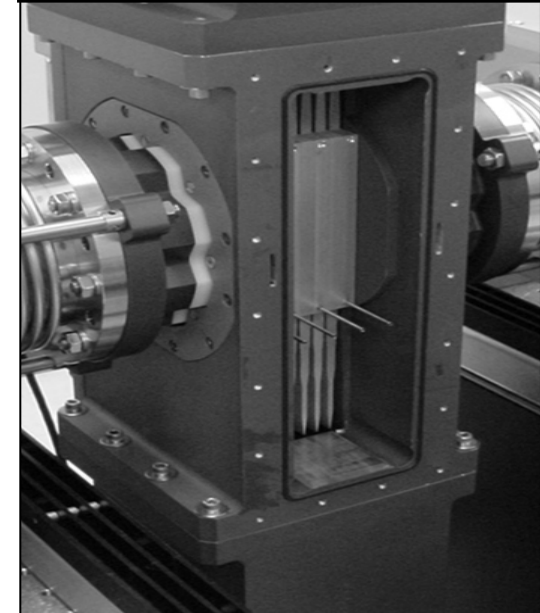
**Diode pump arrays**



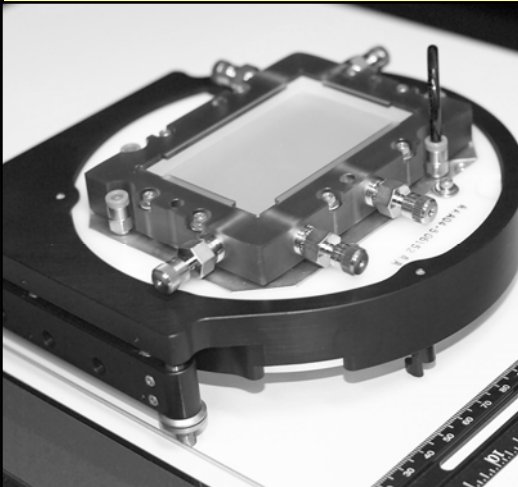
**Solid-state amplifier**



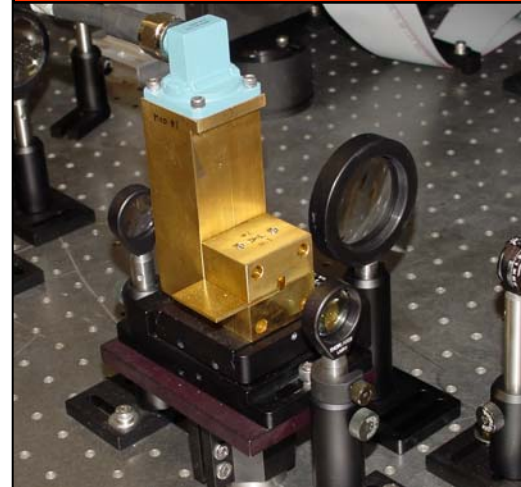
**Helium gas cooling**



**Frequency Converter**



**Broadband Front End**

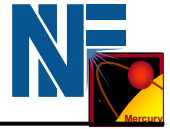


**Adaptive Optic**



# Mercury's front end design uses fiber technology for a stable and robust system meeting NIF (IFE) requirements

---



## Energy stability and beam quality required for reliable ignition pulses

- 500 +/- 2.5 mJ @ 10 Hz
- 10,000:1 signal to noise
- Beam quality: < 1.5X diffraction limited

## Temporal shaping required to compensate for gain saturation and ignition waveforms

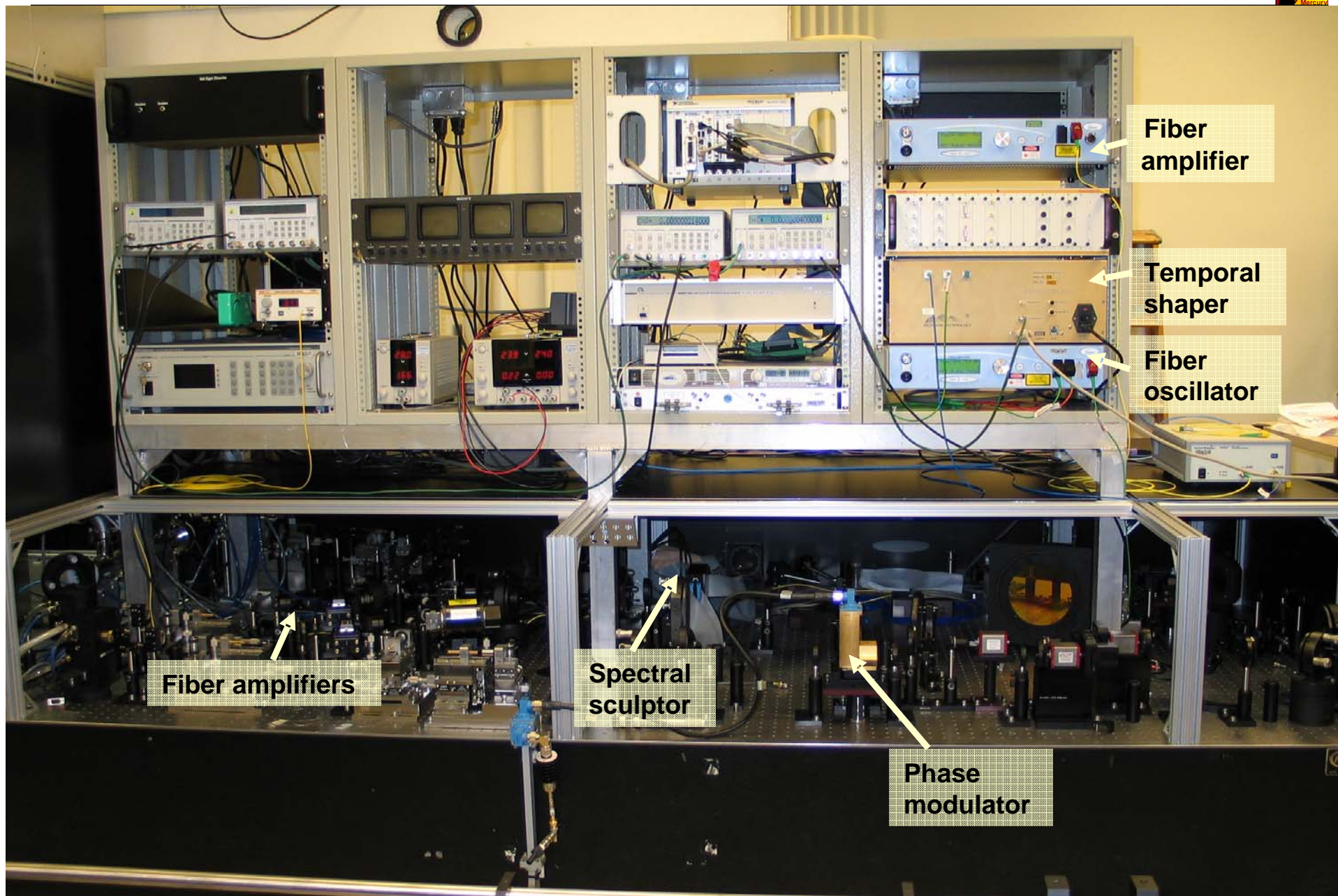
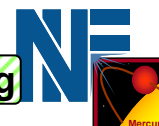
- < 5% amplitude fluctuations
- > 250 ps jitter
- 20:1 contrast

## Spectral bandwidth required for beam smoothing on target

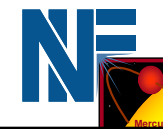
- 3 GHz stability
- >150 GHz bandwidth
- 100:1 contrast



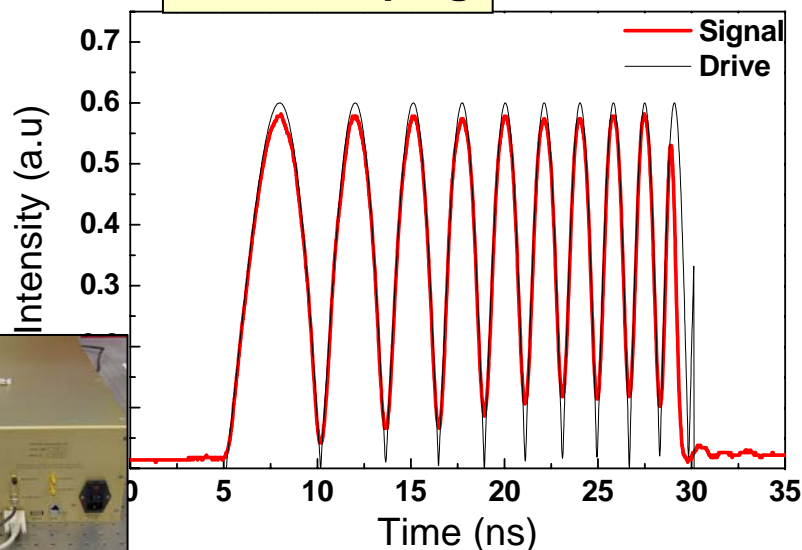
Fiber osc — Temporal shp — Fiber amp — RF bandwidth — Spectral sculpt — Fiber amp — S-FAP Ring



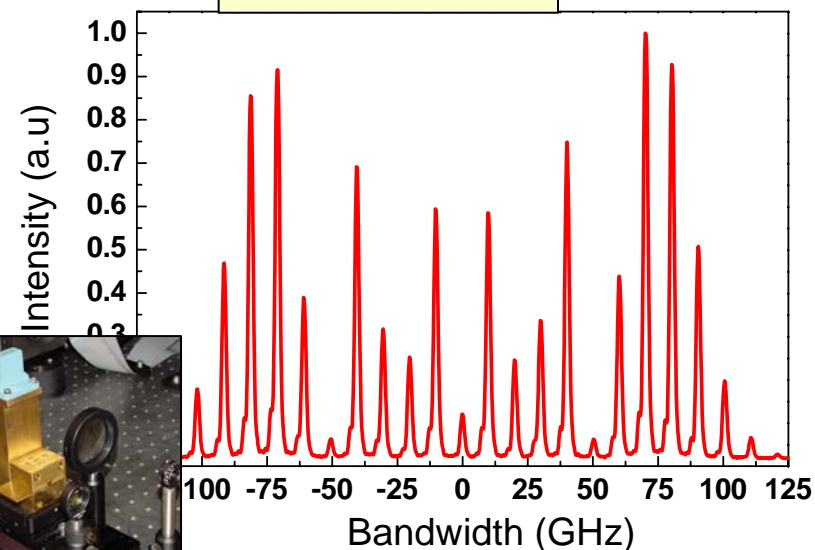
# Front end development progress



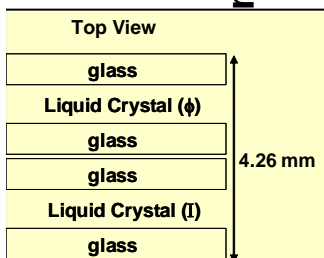
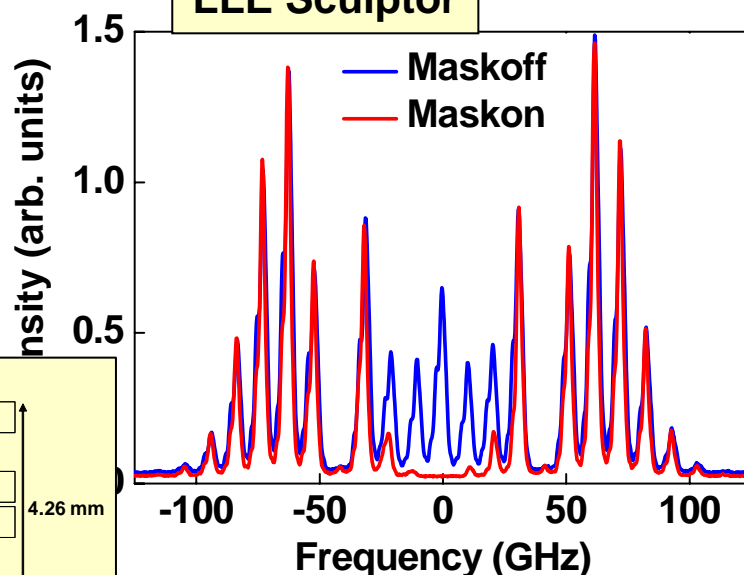
## Pulse shaping



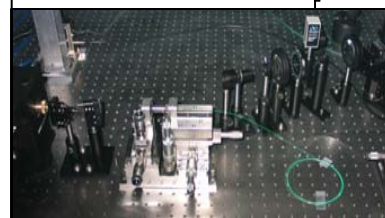
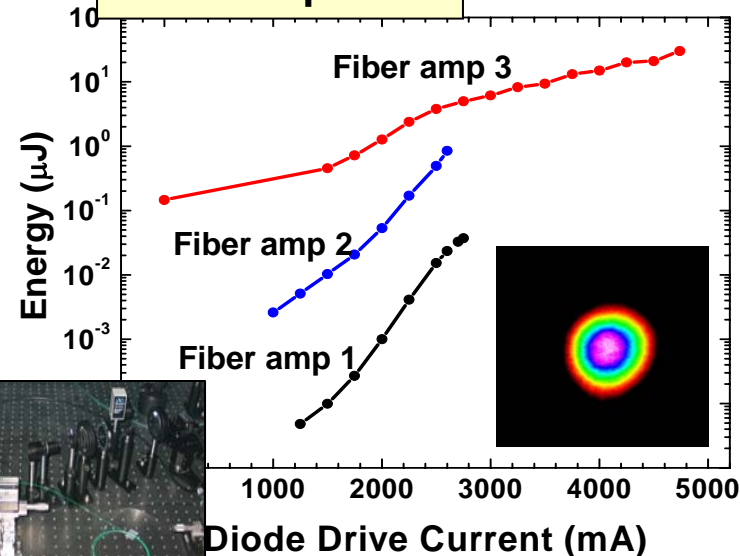
## LLE Modulator



## LLE Sculptor

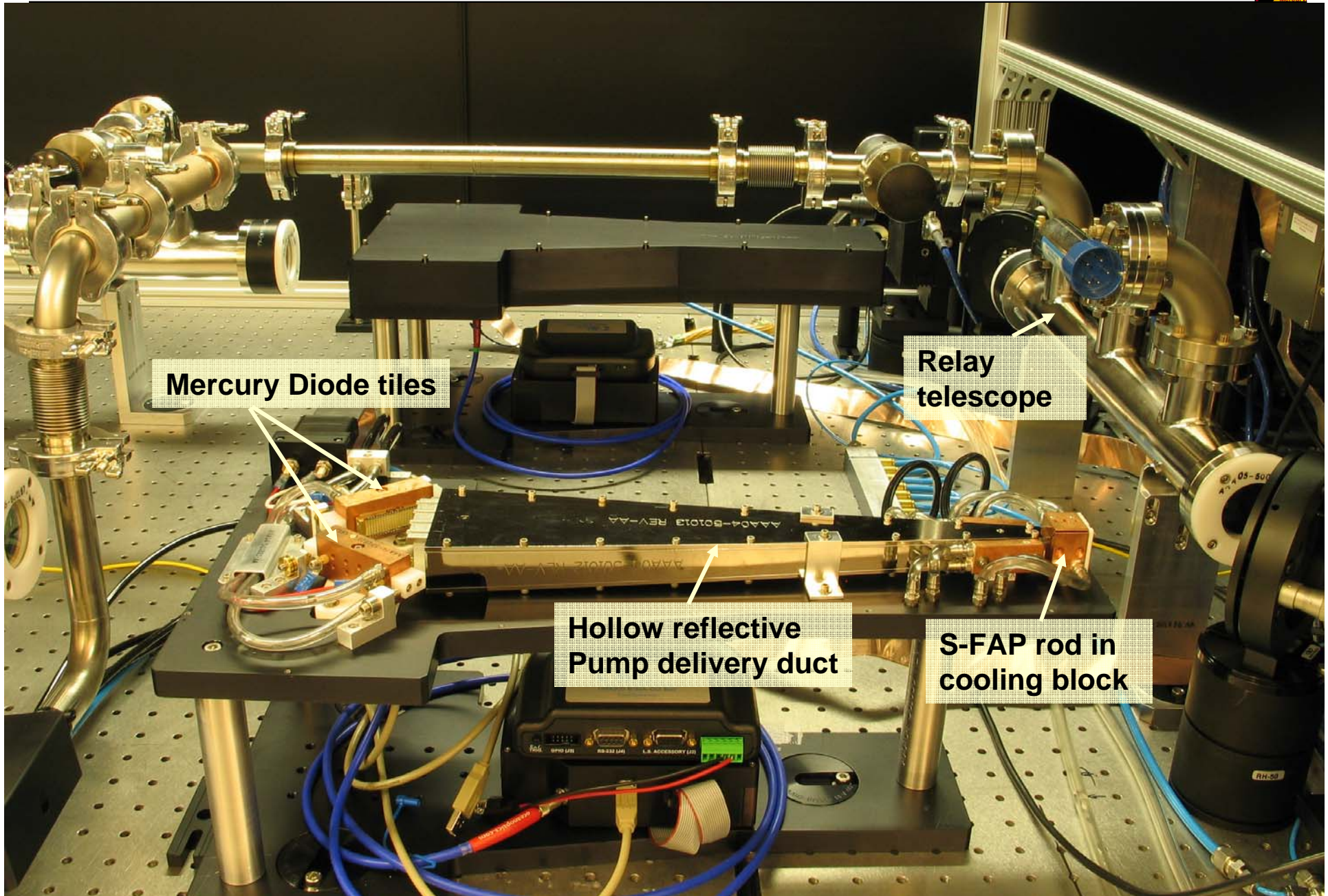
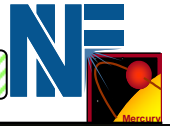


## Fiber Amplifiers





Fiber osc — Temporal shp — Fiber amp — RF bandwidth — Spectral sculpt — Fiber amp — S-FAP Ring



Mercury Diode tiles

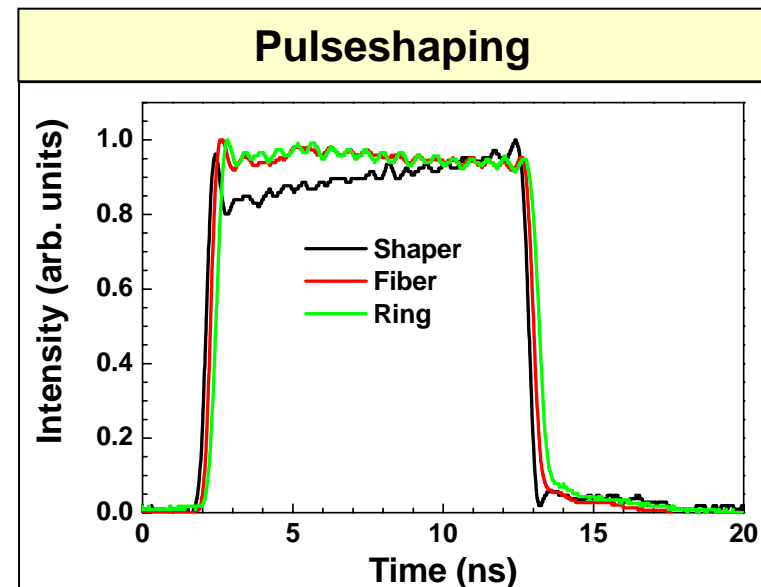
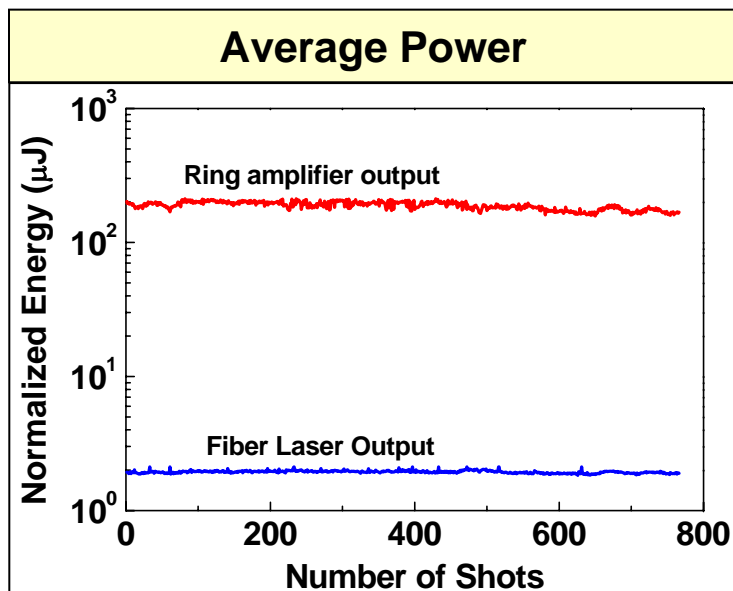
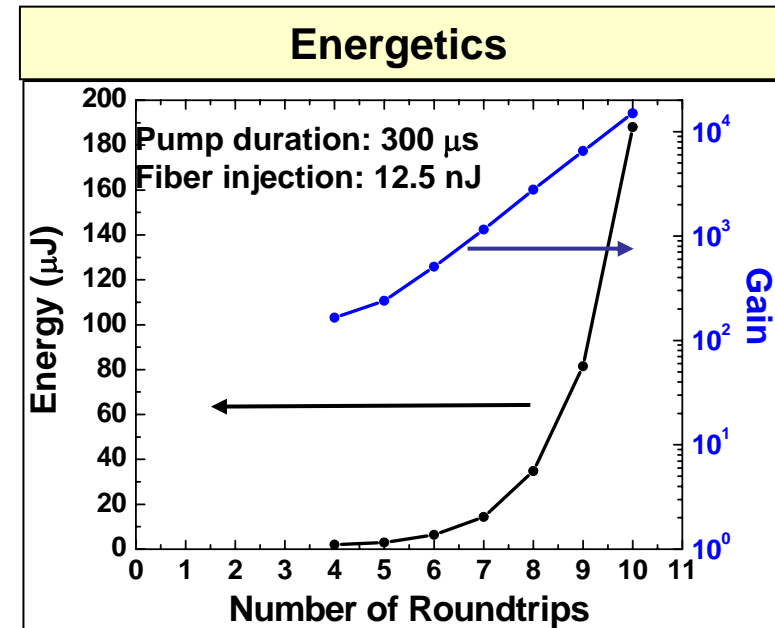
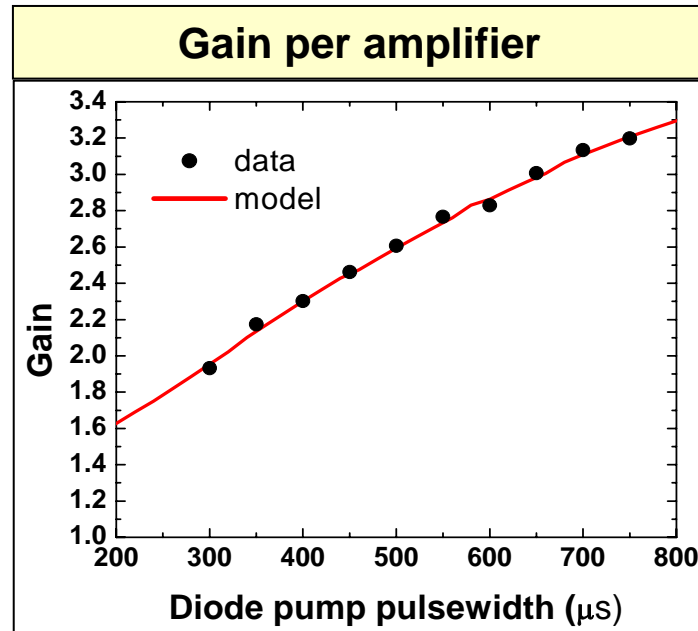
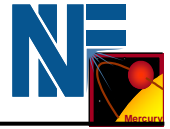
Relay telescope

Hollow reflective Pump delivery duct

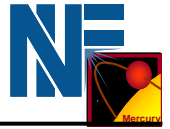
S-FAP rod in cooling block



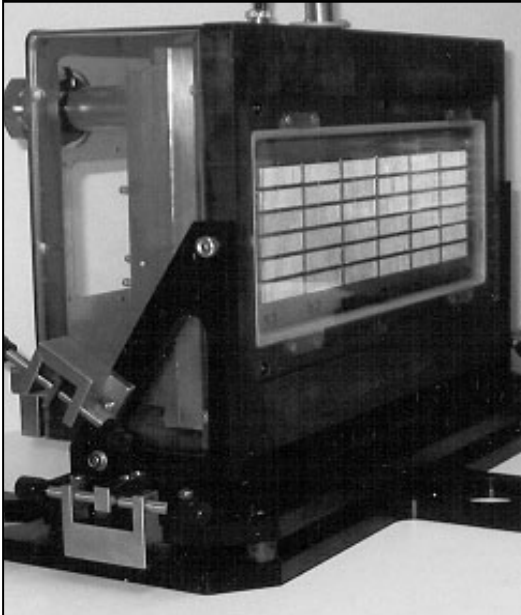
# Initial ring amplifier data show the small signal gain matching the model and successful multipass pulsed amplification



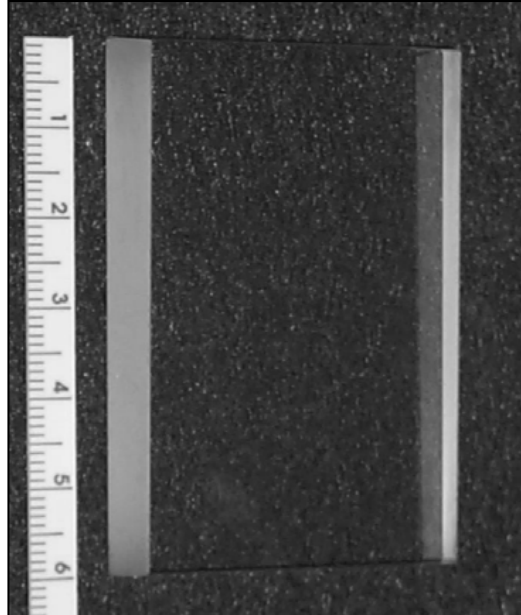
# A new adaptive optic has been installed



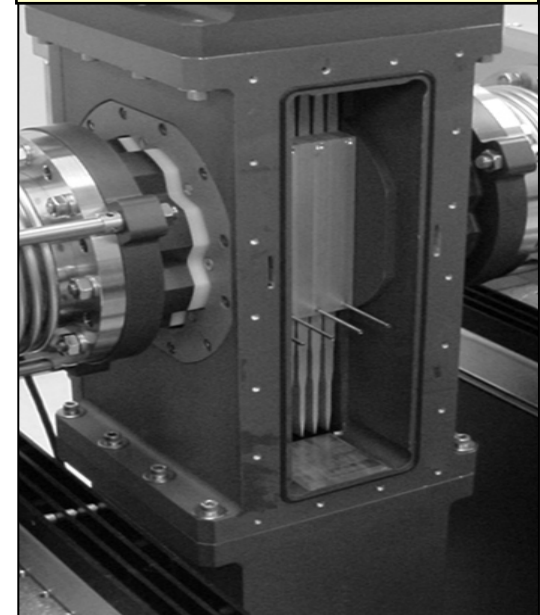
**Diode pump arrays**



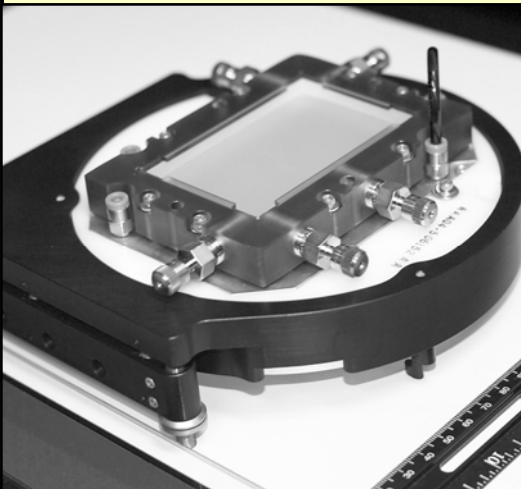
**Solid-state amplifier**



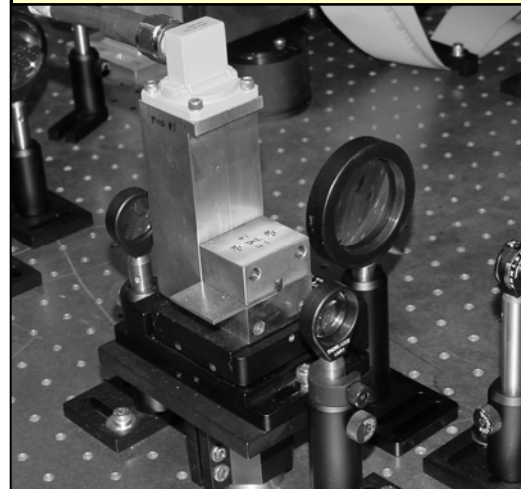
**Helium gas cooling**



**Frequency Converter**



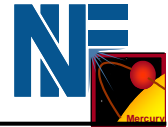
**Broadband Front End**



**Adaptive Optic**

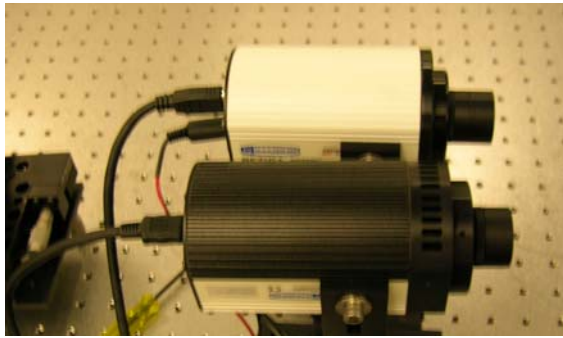


# Our adaptive optic system passed initial inspection and tests



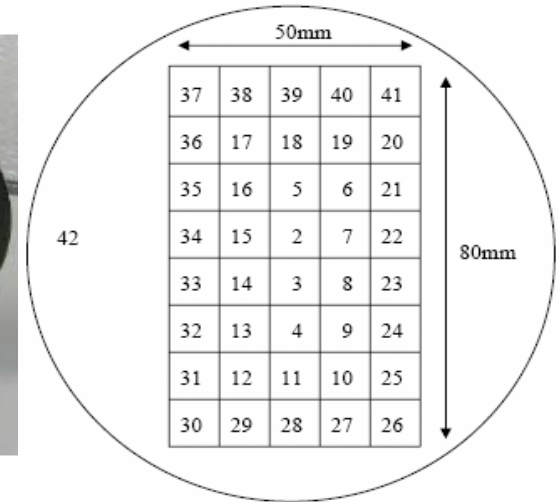
## Wavefront Sensor

Uses a four-wave lateral shearing interferometer for high resolution and simple alignment



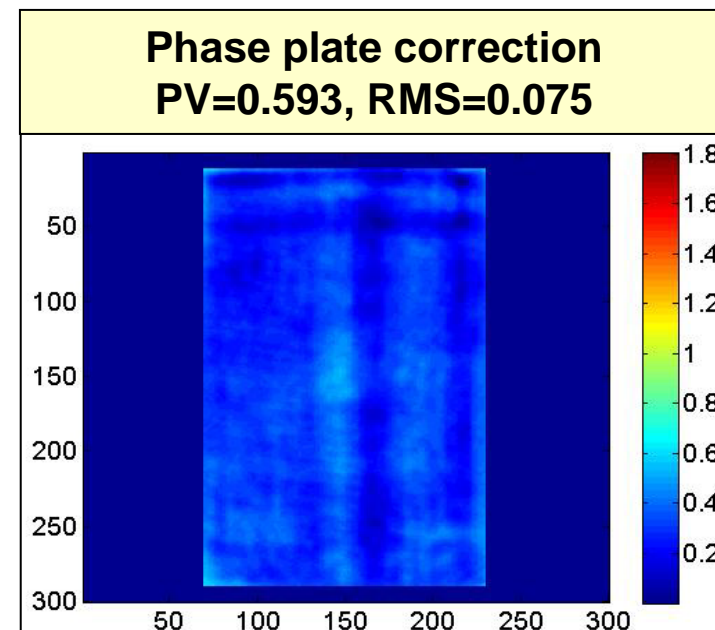
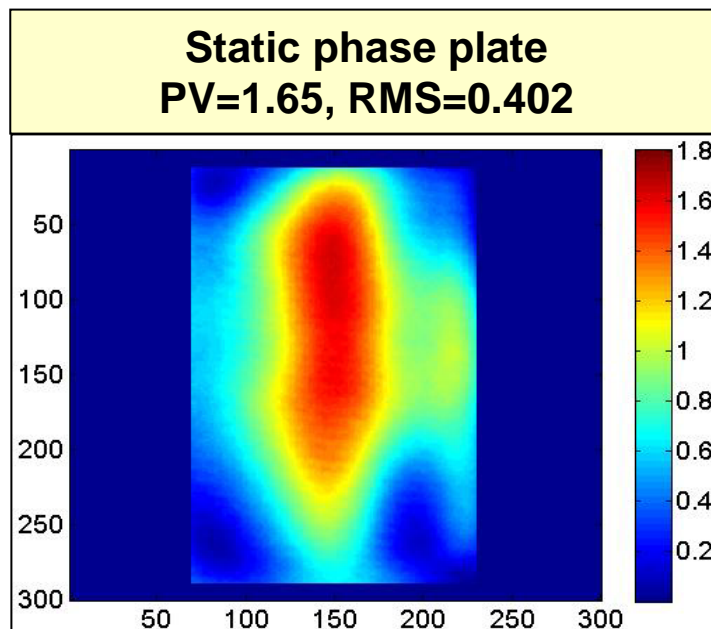
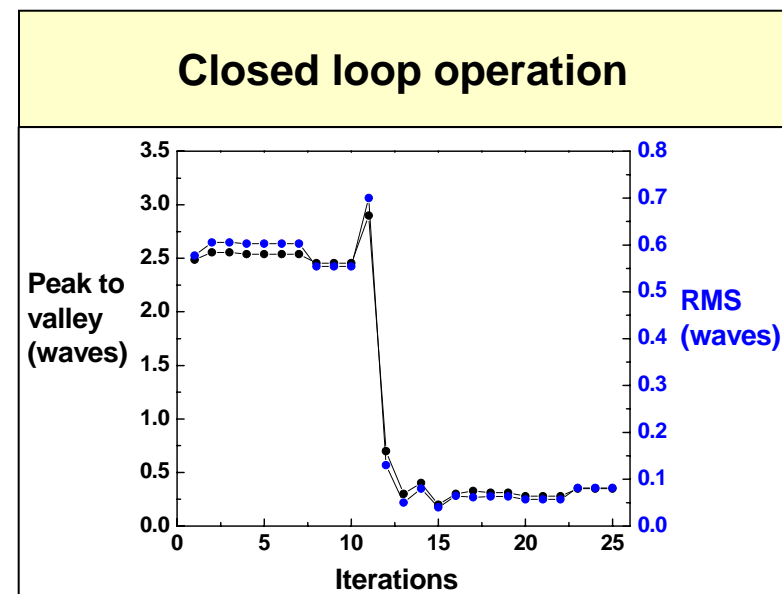
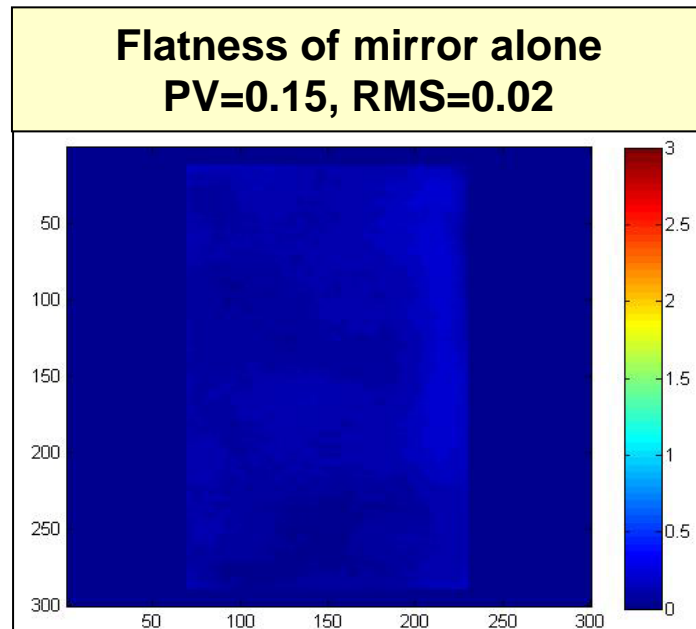
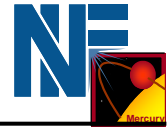
	Specification	Test Result
Resolution	256x256	300x300
Sensitivity (waves)	0.05	0.05
Dynamic Range (waves)	0.05-150	0.05-150
Repetition rate (Hz)	12	12

## Adaptive Optic



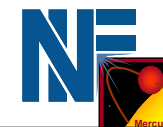
	Spec	Test
Active aperture	45 x 75 mm	OK
Maximum Stroke (w/o defocus)	> 3 waves	OK
Flatness (w/o Defocus)	< 1 wave	0.9 waves
Transmission	< 0.2 %	0.065 %
Damage Threshold	>3 GW/cm <sup>2</sup>	3.3 GW/cm <sup>2</sup>

# A static phase plate was used to test the adaptive optic





## We are successfully meeting our performance goals



Components	Goal		Present	End FY06
Amplifier slabs (4x6 cm)	14	✓	14	28
Diode tiles (120 W/bar)	288	✓	360	360
Amplifiers	2	✓	2	2
- Cooling uniformity (rms)	<1%	✓	0.12%	0.12%
2 $\omega$ Conversion crystals	2	✓	3	3
3 $\omega$ Conversion crystals	3		1	3
Wavefront control	DM		Commissioned	Online
Laser Performance				
Energy (J)	100		65	100
Rep-rate (Hz)	10	✓	10	10
Efficiency (%)	10		6.5	10
Diode reliability (shots)	10 <sup>8</sup>	✓	10 <sup>9</sup>	10 <sup>9</sup>
Beam quality (xDL)	5		4 @ 65J (80%)	5
Pulse-shaping (ns)	3-10	✓	3-15	3-15
Bandwidth (GHz @ 1 $\omega$ )	>150		Offline	250 GHz
Conversion	2 $\omega$ /3 $\omega$		2 $\omega$	2 $\omega$ /3 $\omega$

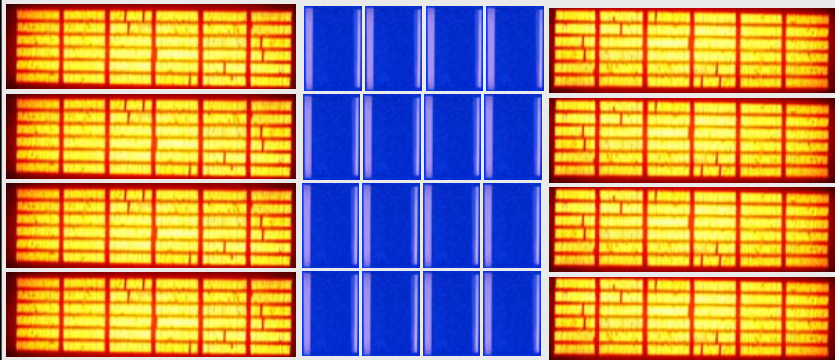
✓ Completed

On schedule

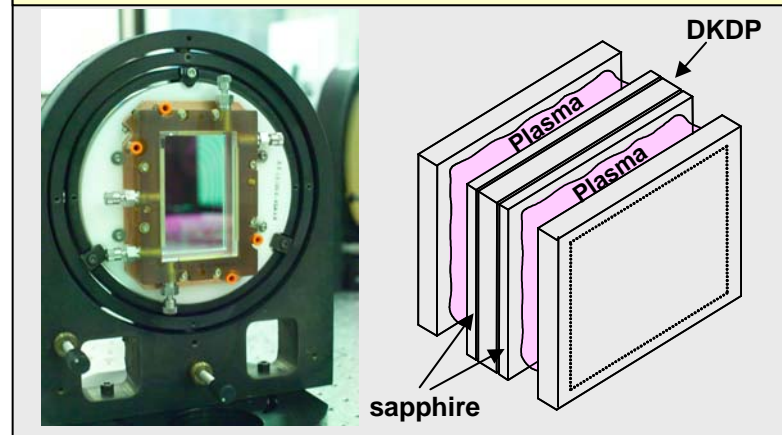
# The DPSSL approach will begin to develop multi-kJ scale components in FY06



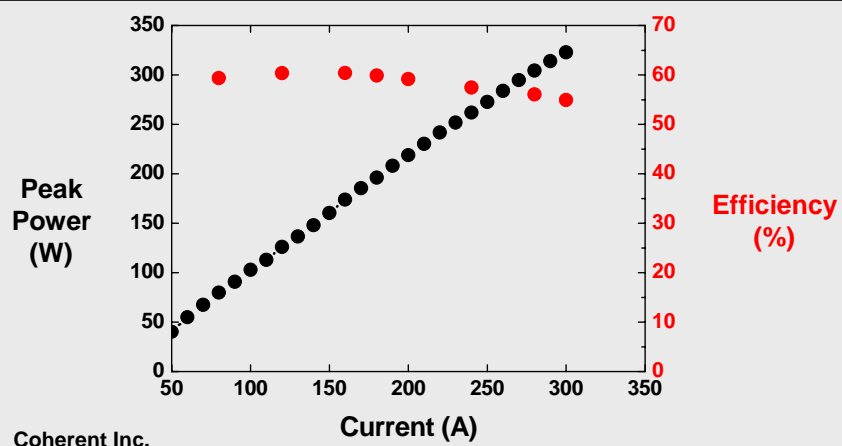
## Beam Bundling



## Optical Switch



## High Performance Diodes



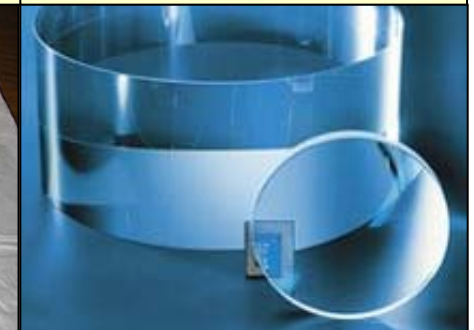
## Large Scale Growth

### 15 cm Nd:GGG



Northrop Grumman

### 40 cm CaF<sub>2</sub>



Schott Technologies

- **Project Overview**

- International/National DPSSL Programs are pushing technology envelopes

- **System Performance**

- Mercury Laser performance goals are on track  
325 W average power at 14 ns for  $> 10^3$  shots (65 J at 5 Hz)  
Beam quality at 80% of energy in a 4X diffraction limited spot

- **Component Performance**

- Pump diode arrays (Commercial prototypes meet specs)
- Crystalline gain media (23 spare slabs in queue)
- Gas cooled amplifiers (Thermal wavefront agrees with model)
- Frequency conversion (YCOB advanced material in production, scalable chilled plate prototype demonstrated)
- Front end (90% complete)
- Adaptive optics (Mirror Commissioned)

- **Next Generation Design Considerations**

- System engineering with statistics in mind
- Out-of-the box thinking to push efficiencies
- Leveraging NIF engineering
- Leveraging subscale facilities for component development/chamber materials/chamber environment

## **Mercury Team**

Kathy Alviso  
Paul Armstrong  
Earl Ault  
Andy Bayramian  
Camille Bibeau  
Ray Beach  
Mike Benapfl  
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Al Erlandson  
Barry Freitas  
Kevin Hood  
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Joe Menapace  
Bill Molander  
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Noel Peterson  
Kathleen Schaffers  
Chris Stolz  
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John Tassano  
Steve Telford  
Everett Utterback

## **Collaborators**

Coherent  
Directed Energy  
Laboratory for Laser Energetics  
Northrop-Grumman  
Onyx Optics  
PHASICS  
Night N (opt) Ltd.  
Crystal Photonics  
Quality Thin Films  
Schott Glass Technologies  
SESO  
Spica  
Zygo

